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Scattering Amplitudes and Gravitational Waves

Snowmass White Paper [Buonanno, Khalil, O'Connell, Roiban, Solon, Zeng; arXiv: [2204.05194](https://arxiv.org/abs/2204.05194)]

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University of California, Los Angeles





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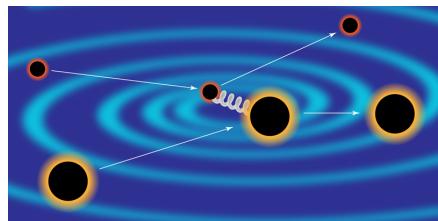
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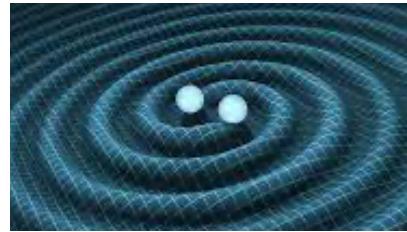
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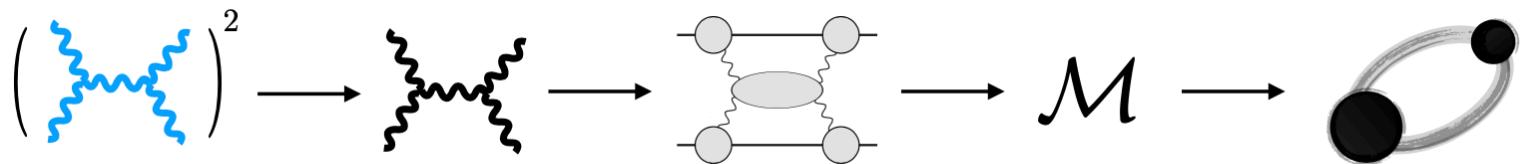
“Black Hole Collider Physics”



Black Hole collision



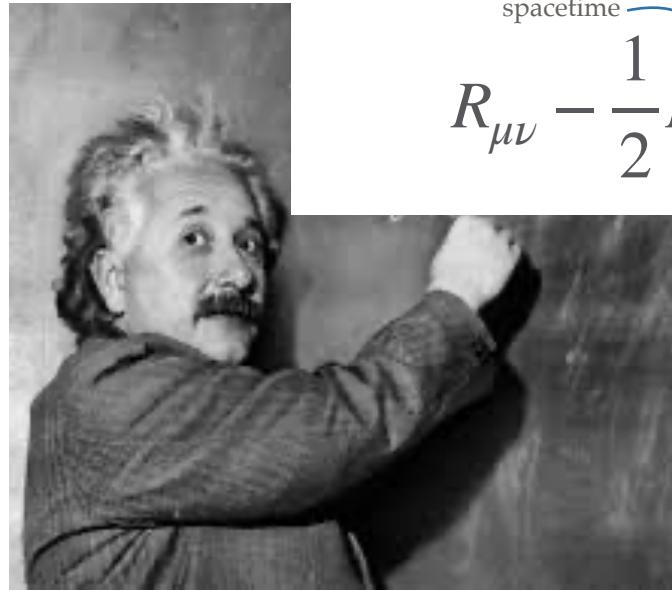
Binary inspiral process



Leverage advanced QFT methods for
state-of-the-art gravitational wave predictions

Gravitational wave physics

One Equation - Extremely Rich Physics



$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

spacetime
matter

SITZUNGSBERICHTE

1918.
VIII.

DER
KÖNIGLICH PREUßISCHEN

AKADEMIE DER WISSENSCHAFTEN.

Gesamtsitzung vom 14. Februar.
Mitteilung vom 31. Januar.

- Expansion of the Universe
- Black Holes
- Gravitational Waves (GWs) **Über Gravitationswellen.**

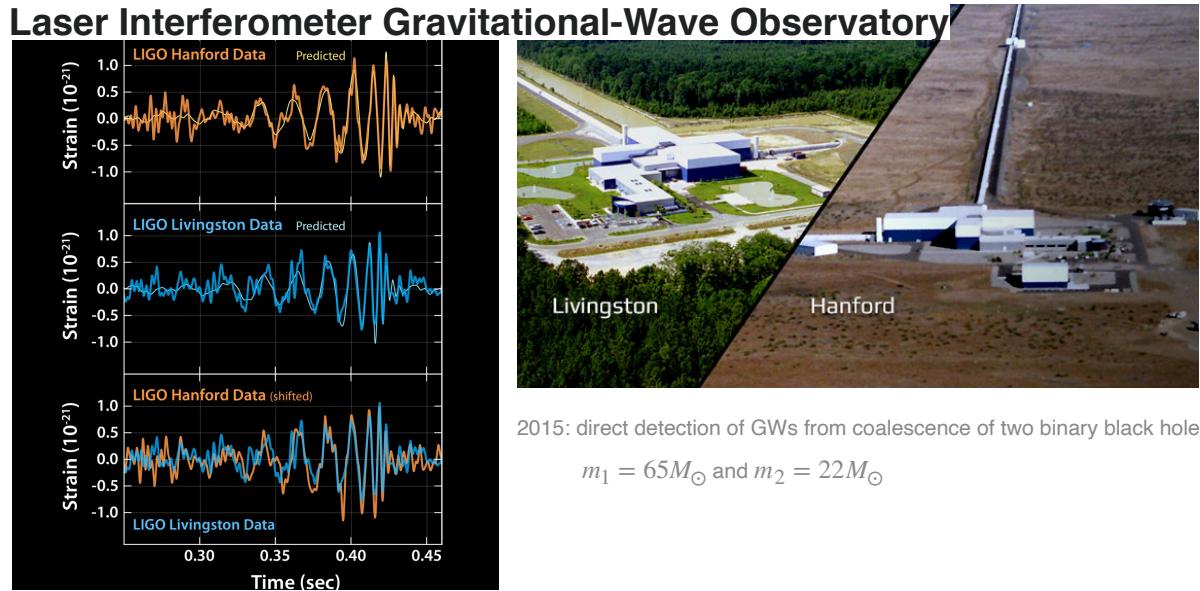
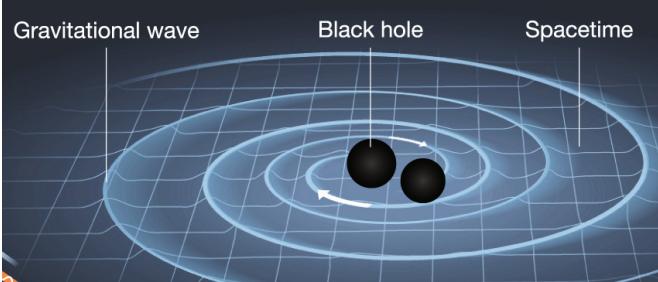
Von A. EINSTEIN.

[Einstein 1915]

Prediction of Gravitational Waves 1918

Gravitational wave physics

Direct detection of Gravitational Waves @ LIGO 2015

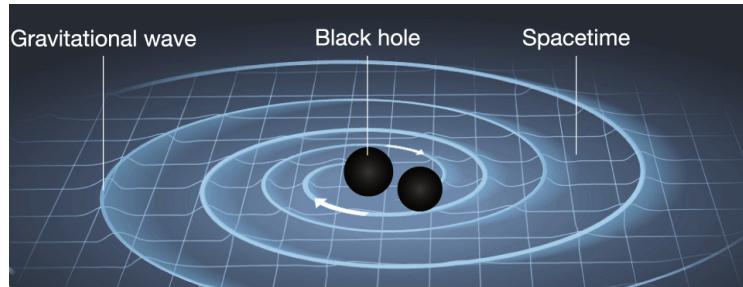


Gravitational waveform determined by General Relativity

Gravitational wave physics

A promising scientific program ahead:

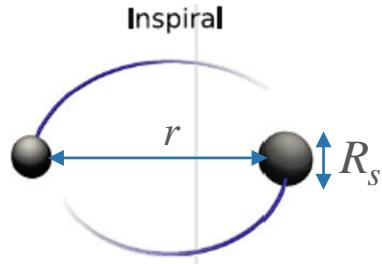
- Study General Relativity in strong field regime
- Constraints on Neutron star equations of state
- Precision tests of GR, modifications of gravity? New Physics? ...



requires theory precision

Gravitational wave physics

Theory of gravitational waveforms:

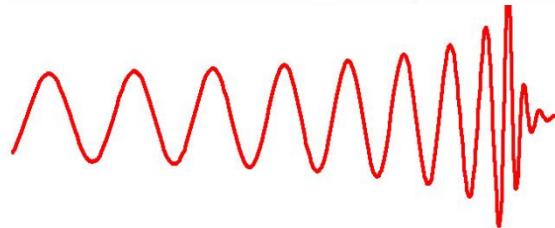
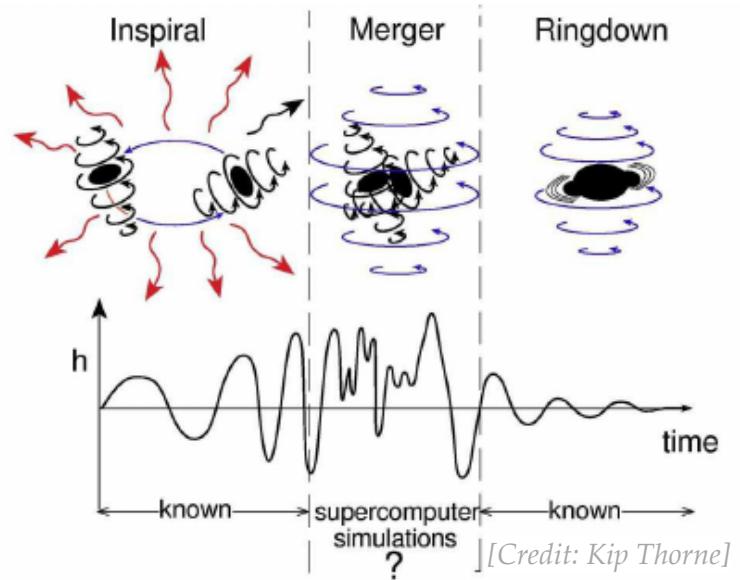


$$\left[\frac{v}{c} \right]^2 \sim \frac{GM}{r} = \frac{R_s}{r} \ll 1$$

virialization

weak field, slow motion — *pert. theory!*
(velocities become large close to merger)

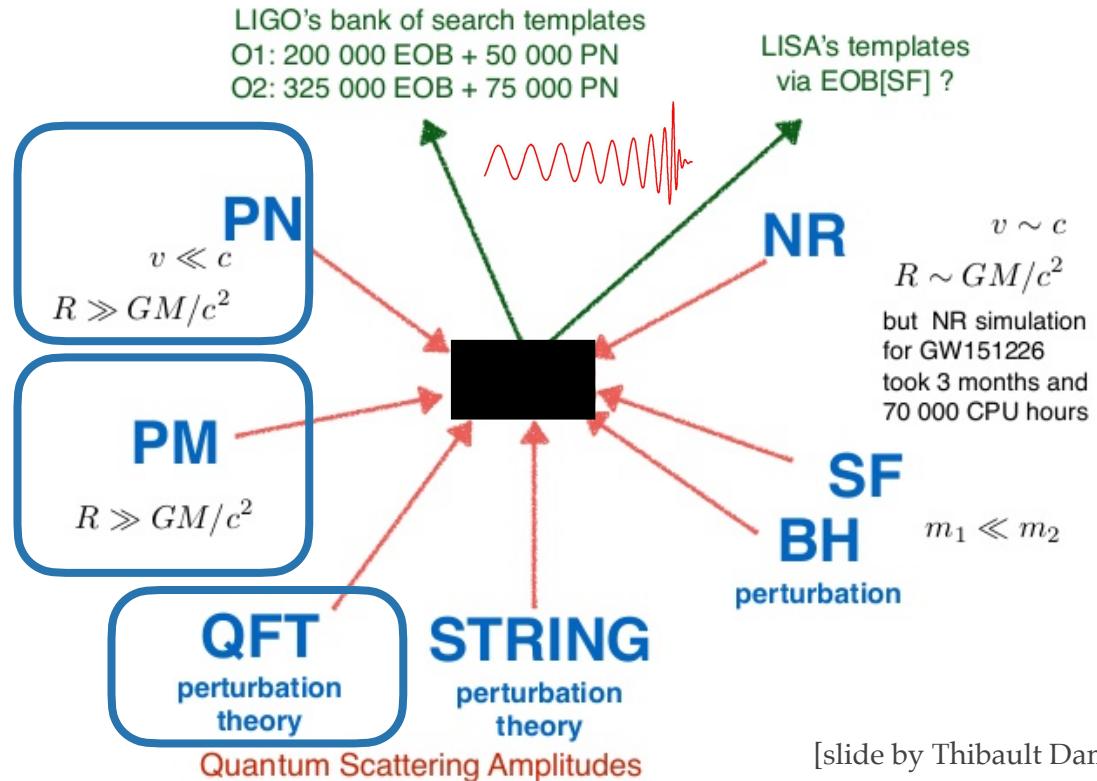
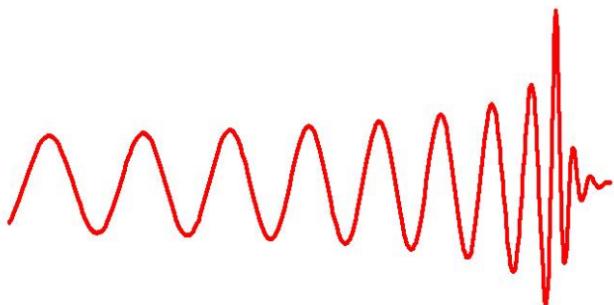
dynamics described by classical
Hamiltonian H



Gravitational wave physics

LIGO theorists:
waveform modeling

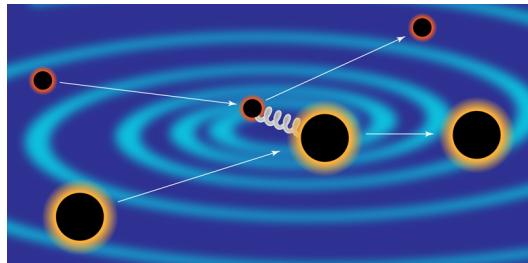
[Buonanno; Damour; ...]



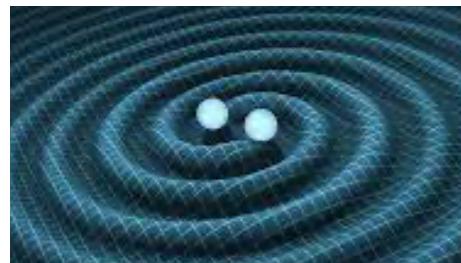
[slide by Thibault Damour]

QFT methods for classical General Relativity

“Black hole collider physics”



Black Hole collision



Binary inspiral process

New source of theory input

Scattering Amplitudes

Encouraged by general relativists:

[Kovacs, Thorne 1978]

g) *The Feynman-Diagram Approach*

Any classical problem can be solved quantum-mechanically; and sometimes the quantum solution is easier than the classical. There is an extensive literature on the Feynman-diagram, quantum-mechanical treatment of gravita-

High-energy gravitational scattering and the general relativistic two-body problem

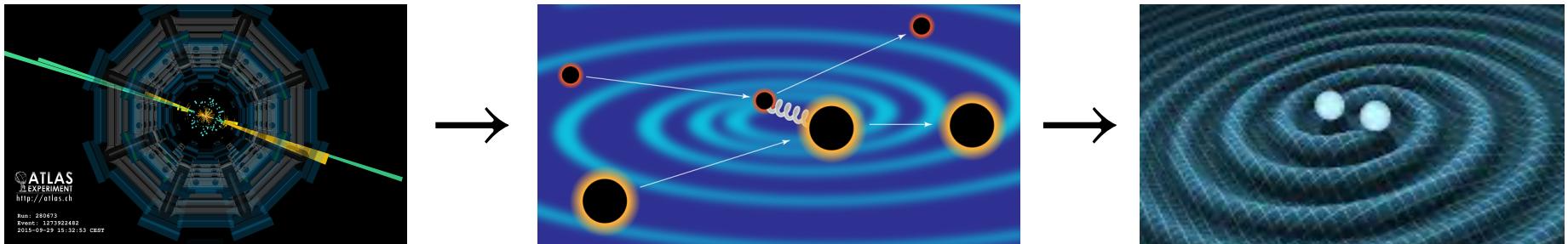
Thibault Damour*

Institut des Hautes Etudes Scientifiques, 35 route de Chartres, 91440 Bures-sur-Yvette, France

(Dated: October 31, 2017)

[...] tum gravitational scattering amplitude of two particles, and we urge amplitude experts to use their novel techniques to compute the 2-loop scattering amplitude of scalar masses, from which one could deduce the third post-Minkowskian effective one-body Hamiltonian

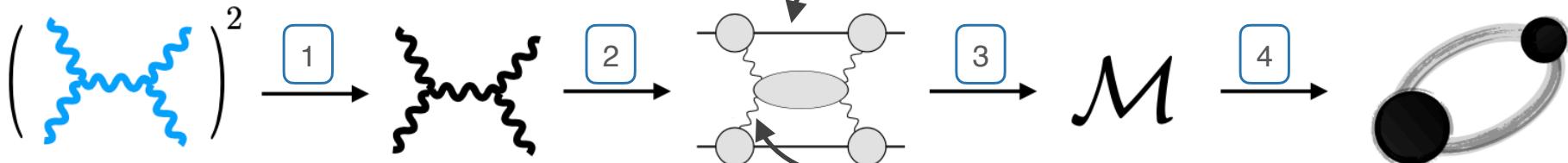
Scattering Amplitudes



What particle theorists bring to the table:

- exploit scale separation: **Effective Field Theory** [Goldberger, Rothstein] [Beneke, Smirnov] [Cheung, Rothstein, Solon]
- export mature collider physics tools (**loop computation**) [Laporta; Tkachov; Chetyrkin; Kotikov; Remiddi, Gehrmann; Henn, Anastasiou, Melnikov, ...]
- focus on **physical**, gauge-invariant **quantities** [Kosower, Maybee, O'Connell] [EH, Parra-Martinez, Ruf, Zeng]
- **Boundary-to-Bound map** between hyperbolic and bound motion [Porto, Kälin]

Scattering Amplitudes



1 double copy

Kawai, Lewellen, Tye, Bern, Carrasco, Johansson, Ochirov, Chiodaroli, Jin, Roiban, Chen, Boucher-Veronneau, Dennen, Davies, Damgaard, Bjerrum-Bohr, Huang, Anastasiou, Borsten, Nagy, ...



2 on-shell methods

Bern, Dixon, Dunbar, Kosower, Morgan, Britto, Cachazo, Feng, Witten, Ellis, Melnikov, Forde, Bader, Carrasco, Johansson, Arkani-Hamed, Y-T Huang, Trnka, Bourjaily ...

3 advanced multiloop integration

Chetyrkin, Tkachov, Laporta, V. Smirnov, A. Smirnov, Kotikov, Bern, Dixon, Kosower, Remiddi, Gehrmann, Chuharev, Tarasov, Lee, Henn, Beneke, Czakon, Ita, Zeng, ...

4 effective field theory

Caswell, Lepage, Braaten, Isgur, Wise, Manohar, Stewart, Donoghue, Beneke, Goldberger, Rothstein, Neill, Vaidya ...

[Credit: Mikhail Solon]

Scattering Amplitudes

Combination of QFT tools led to new state-of-the-art results in GR
(Post-Minkowskian [PM] expansion: Perturbation theory in Newton's constant G only)

$$G^1 (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) \quad 1 \text{ PM}$$

$$G^2 (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) \quad 2 \text{ PM}$$

$$G^3 (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) \quad 3 \text{ PM}$$

$$G^4 (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) \quad 4 \text{ PM}$$

$$G^5 (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) \quad 5 \text{ PM ?}$$

1-loop

[Westpfahl 1985]

2-loop

[Bern, Cheung, Roiban, [Kälin, Liu, Porto 2021]
Shen, Solon, Zeng 2019] (conservative)

3-loop

[EH, Parra-Martinez, Zeng 2021]
[Vanhove et al. 2021] +(radiation)

4-loop

[Bern et al. 2021] [Porto et al. 2021] (conservative)

Relativity built into QFT approach – all orders in velocity results

Scattering Amplitudes

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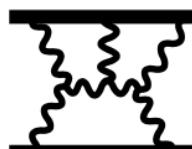
$$G^3 (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) \quad 3 \text{ PM}$$

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$$G^5 (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) \quad 5 \text{ PM}$$



Bini, Damour, Geralico 19, 20
Damour 19, 20
Blümlein, Maier, Marquard, Schäfer 20
Di Vecchia, Heissenberg, Russo, Veneziano 20
EH, Parra-Martinez, Ruf, Zeng 21

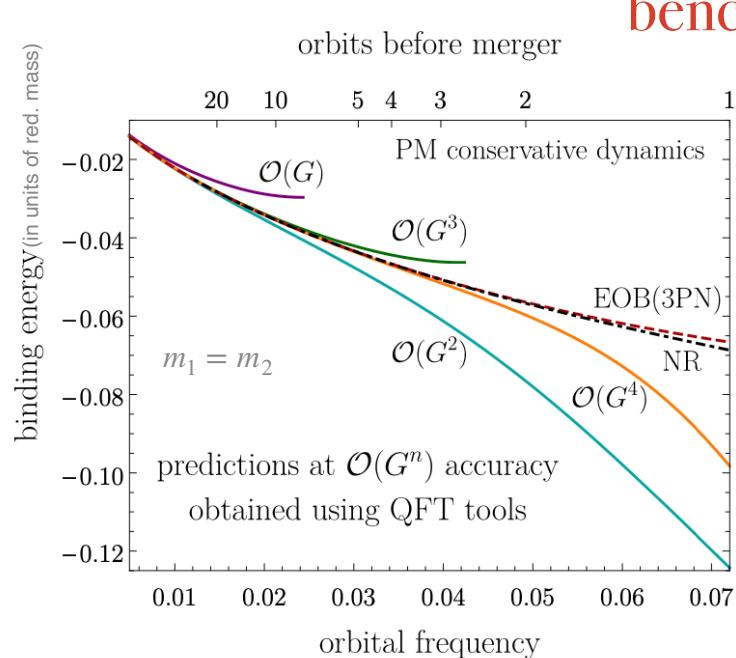


Bini, Damour, Geralico 21
Blümlein, Maier, Marquard, Schäfer 21
Bern, Parra-Martinez, Roiban, Ruf,
Shen, Solon, Zeng 21
Dlapa, Kälin, Liu, Porto 21

Relativity built into QFT approach – all orders in velocity results

Scattering Amplitudes

[Buonanno, Khalil, O'Connell, Roiban, Solon, Zeng; arXiv:[2204.05194](https://arxiv.org/abs/2204.05194)]



[Antonelli, Buonanno, Steinhoff, van de Meent, Vines arXiv: [1901.07102](https://arxiv.org/abs/1901.07102)]

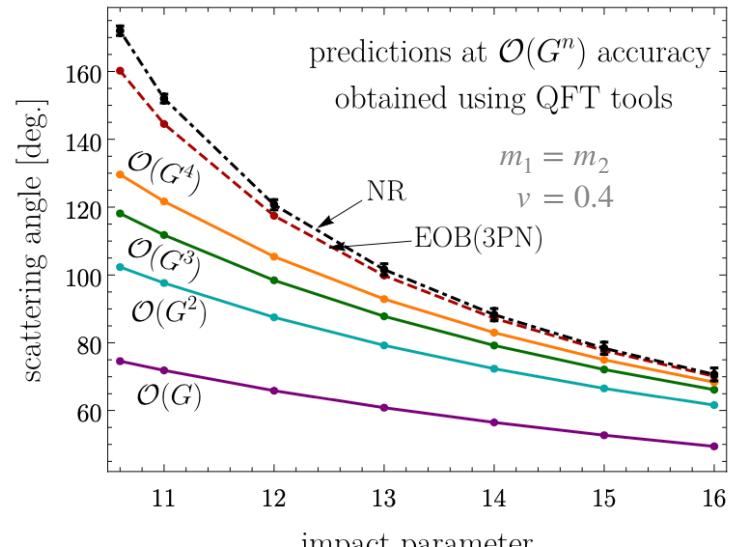
bound



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Amplitudes and Gravitational Waves

benchmark quantities



[Khalil, Buonanno, Steinhoff, Vines, arXiv: [2204.05047](https://arxiv.org/abs/2204.05047)]

unbound

Current status / Outlook

spin and tidal effects



Guevara, Ochirov, Vines, Huang, Levi, Bern, Luna, Roiban, Shen, Zeng, Kosmopoulos, Jakobsen, Mogull, Cheung, Shah, Haddad, Helset, Chia, Charalambous, Dubovsky, Ivanov, Hui, Joyce, Penco, Santoni, Solomon, De Luca, Pani ...

radiation



Damour, Veneziano, Di Vecchia, Heissenberg, Russo, Herrmann, Parra-Martinez, Ruf, Zeng, Riva, Vernizzi, Cristofoli, Kosower, O'Connell, Jakobsen, Mogull, Plefka, Steinhoff, Bjerrum-Bohr, Damgaard, Plante, Vanhove, Shen, Manohar, Ridgway ...

classical double copy

Luna, Monteiro, O'Connell, White, Nicholson, Ridgway, Wise, Carrillo-Gonzalez, Penco, Trodden, Kol, Adamo, Casali, Mason, Nekovar, Ochirov, Goldberger, Ochirov, Plefka, Shi, Shen, Arkani-Hamed, Huang, Bjerrum-Bohr, Donoghue, Vanhove, Gonzo ...

theoretical structures



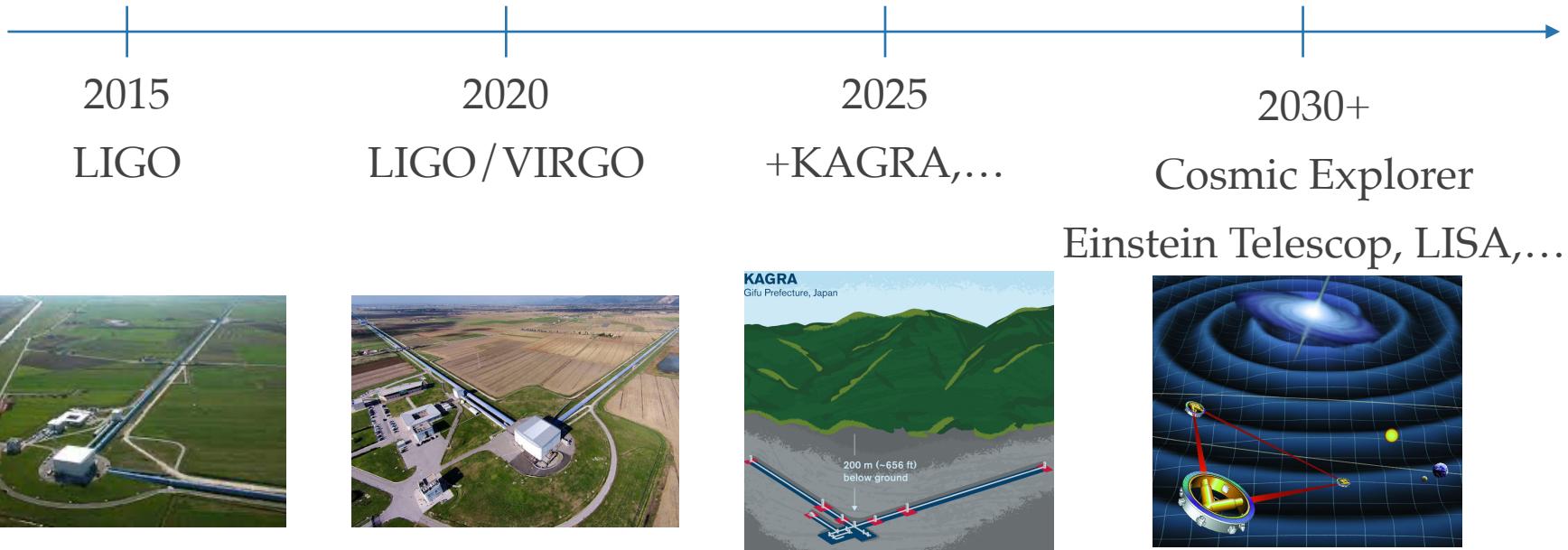
Cachazo, Guevarra, Ita, Ruf, Hermann, Luna, Zeng, Jakobsen, Mogull, Plefka, Steinhoff, Bjerrum-Bohr, Damgaard, Vanhove, Kol, Telem, O'Connell, Cristofoli, Gonzo, Kosower, Britto, Brandhuber, Johansson, Travaglini, Carrillo-Gonzalez, de Rham, Tolley, Jones, Gralla ...

new methods

Kosower, Maybee, O'Connell, Bini, Damour, Geralico, Bjerrum-Bohr, Damgaard, Festuccia, Plante, Vanhove, Kalin, Porto, Parra-Martinez, Ruf, Zeng, Brandhuber, Chen, Johansson, Travaglini, Wen, Edison, Levi, Di Vecchia, Russo, Heissenberg, Veneziano, Chiodaroli, Johansson ...

BH vs. Neutron Star

Outlook



The future of particle physics at interface with GW science is bright!

Outlook

[Buonanno, Khalil, O'Connell, Roiban, Solon, Zeng; arXiv: [2204.05194](https://arxiv.org/abs/2204.05194)]

Snowmass White Paper: Gravitational Waves and Scattering Amplitudes

Alessandra Buonanno,^{1,2} Mohammed Khalil,^{1,2} Donal O'Connell,³ Radu Roiban,⁴
Mikhail P. Solon,⁵ Mao Zeng³

lem in general relativity. Theorists developing waveform models for the LIGO-Virgo-KAGRA (LVK) collaboration [1–3] have performed initial studies [4, 5] of these early results, see Figure 1, and have strongly encouraged further developments. Indeed, if these calculations are pushed to higher orders, and are extended to include all physical effects (i.e., spins and tides), they can be used, in combination with other analytic methods [6–11]¹ and with numerical-relativity (NR) simulations [16–19], to provide highly accurate waveform models of binary systems composed of black holes and/or neutron stars. Another aspect of this program that has drawn significant interest from

Relevant Snowmass White Papers

[Buonanno, Khalil, O'Connell, Roiban, Solon, Zeng; arXiv: [2204.05194](#)]

Snowmass White Paper: Gravitational Waves and Scattering Amplitudes

[Adamo, Carrasco, Carrillo-Gonzalez, Chiodaroli, Elvang, et al.; arXiv: [2204.06547](#)] **Snowmass White Paper: the Double Copy and its Applications**

[Goldberger; arXiv: [2206.14249](#)]

Snowmass White Paper: Effective field theories of gravity and compact binary dynamics

Thank you!